

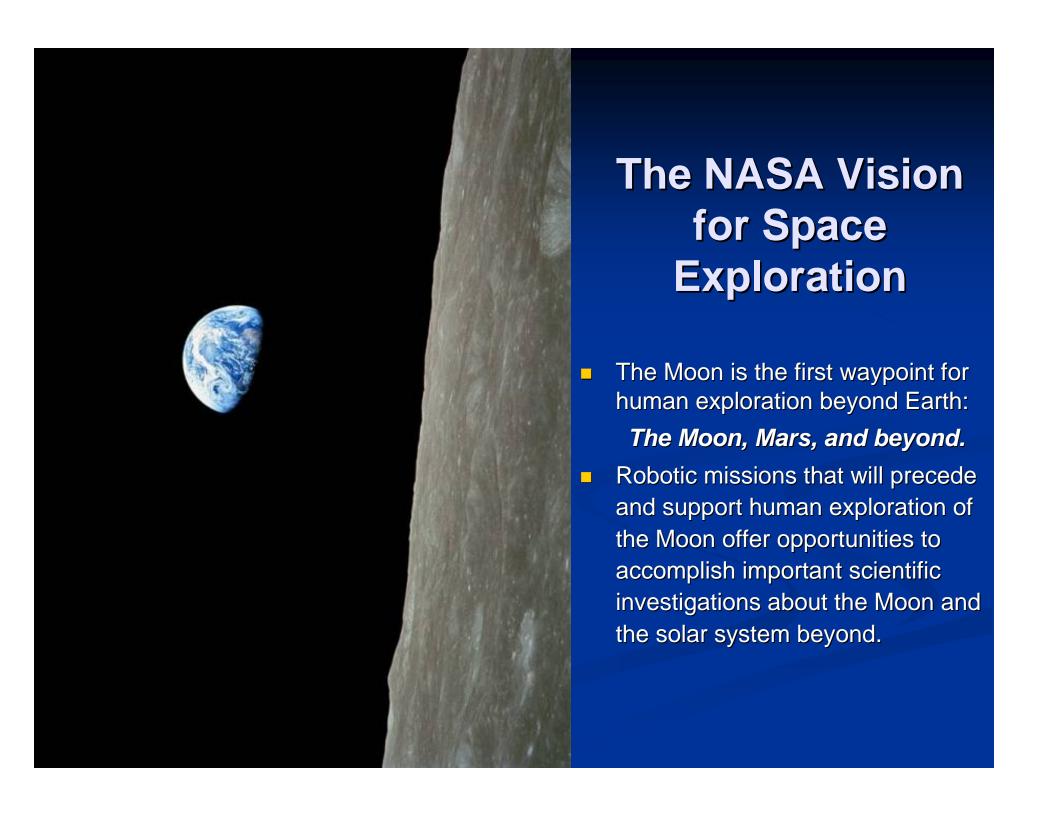
3 SPACE EXPLORATION CONFERENCE & EXHIBIT

Robotic Missions to the Moon: Science & Exploration

Presenter: Dr. Carle Pieters

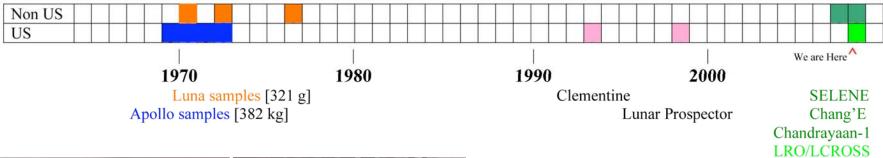
Title: Professor, Brown University

Date: February 27, 2008



Lunar Exploration Timeline: Return after a long Drought













- The Apollo/Luna samples brought new and fundamental understanding of planetary evolution (and the Earth-Moon system).
- After decades of neglect, two very small missions were sent to the Moon. The small pulse of new data sparked several paradigm shifts.
- A fleet of sophisticated modern sensors are now at last exploring the Moon.

Post-Apollo Scientific Hypotheses

The context for understanding the origin and evolution of the Moon



Explains the origin of the Moon as being assembled from debris after the impact of a Mars-sized object with the early Earth.

• The lunar magma ocean hypothesis

Governs understanding of the formation of lunar rocks following lunar formation, and suggests that the outer portions of the Moon (several hundred kilometers in depth) were entirely molten. Differentiation of the vast magma body, a magma ocean, resulted in the formation of the earliest crust and mantle and produced the rocks observed today:

• The terminal cataclysm (Late Heavy Bombardment) hypothesis

Concerns the timing of the impact flux in the 600 Ma after lunar formation. It proposes that the largest craters observed on the Moon (vast multi-ringed impact basins) were formed in a brief pulse of impacts of large objects near 4 Ga ago, well after impact-causing debris left over from solar system formation had died away. [An alternate hypothesis is that the rate of impacts to the Moon and Earth declined with time and no cataclysm occurred.]

Time Since Earth Formation (b,y.)
2.0
1.5
1.0
0.5
0

Tu/2 - 10 m.y.

Tu/2 - 22 m.y.

Late Heavy
Bombardment
Accretion
2.5
3.0
3.5
4.0
4.5

Paradigm Shifts from two small missions

- The enormous South-Pole Aitken basin dominates the feldspathic farside of the Moon.
 - Largest and oldest lunar basin
 - Minor basalt fill
 - Iron-rich interior (lower crust/mantle)
- Heat producing elements were concentrated on the lunar nearside (Apollo sites) early in lunar history.
- The poles are unusual environments and may accumulate volatiles.

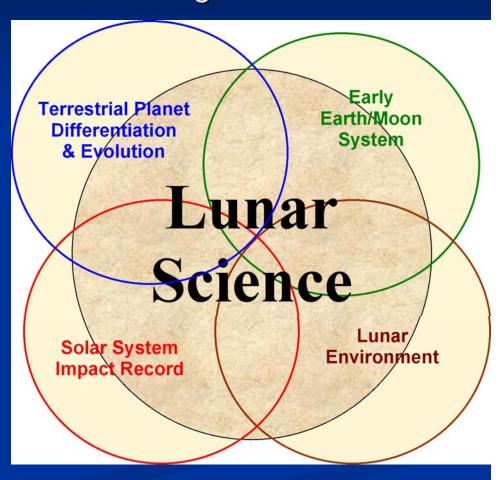
Clementine Topography Albedo Color Composite Lunar Prospector Iron **Thorium**

Polar H

Why the Moon?

- The Moon is a witness to 4.5 billion years of solar system history.
- The Moon presents a record of planetary geologic processes in the purest form
 - Early crust evolution
 - Differentiation
 - Impact craters
 - Volcanic processes
 - Regolith processes and early Sun
- The Moon provides accessible unique environments
 - Polar regions
 - Exosphere (atmosphere)
 - Stable Platform

Overarching Science Themes



See NRC/NAS 2007 Report:

http://books.nap.edu/catalog.php?record_id=11954

International Lunar Exploration

	SELENE/KAGUYA [JAXA]	Chang'E [CNSA]	Chandrayaan1 [ISRO]	LRO [NASA]
Launch	2007	2007	2008	2008
Orbit	100 km polar	200 km polar	100 km polar	50 km polar
	circular	circular	circular	circular
Objectives	Study lunar origin	Surface	Simultaneous	Improve
	and evolution;	structure,	composition	geodetic net;
	develop technology	topography,	and terrain	evaluate polar
	for future lunar	composition;	mapping;	areas; study
	exploration	particle	demonstrate	radiation
		environment	impact probe	environment
Payload	TC, MI, SP, relay	4-band	TMC, HySI,	LOLA, LROC,
	satellites, X-ray, g-	micro-wave,	LLRI, HEX,	LAMP, <i>LEND</i> ,
	ray; laser altimeter;	IIM, X-ray,	Impact probe +	CRaTER,
	radar sounder,	gamma-ray,	C1XS, SARA,	Radiometer,
	magnetometer,	WA stereo,	SIR2,	[miniRF]
	plasma imager	energetic	miniSAR, M3,	LCROSS
		ions, laser	RADOM	
		altimeter		











Modern Remote Sensing Instruments at the Moon



Optical Sensors
Unique
*best of class:

Instrument Type	SELENE	ChangÕE	Chandrayaan-1	LRO
	[2007 JAXA]	[2007 CNSA]	[2008 ISRO]	[2008 NASA]
Stereo imaging	TC	X	TMC*	
<1 m camera				LROC
Spectral camera	MI			(LROC)
VIS+ image		IIM	HySI	
spectrometer				
Near-IR point	SP		SIR2	
spectrometer				
Near-IR ⁺ image			M3	
spectrometer				
Laser altimeter	LALT	X	LLRI	LOLA*
X-ray Fl	XRS	X	C1XS, LEX	
γ-ray	GRS	X	HEX	
Neutron detector				LEND
Plasma/ions	CPS, PACE	X, X	RADOM, SARA	CRaTER
UV Imager				LAMP
Gravity relay	RSAT			
SAR			MiniSAR	MiniRF
Microwave	LRS	Sounder [4λ]		
sounder				
Radiometer				Diviner
Magnetometer	MAG/ER			
Other	HDTV		Impact Probe	LCROSS*
Planned	Lander-	Lander;	Lander	Lunar base
Follow-on	rover	Sample		
		return		
Planned	Lander-	Sample		

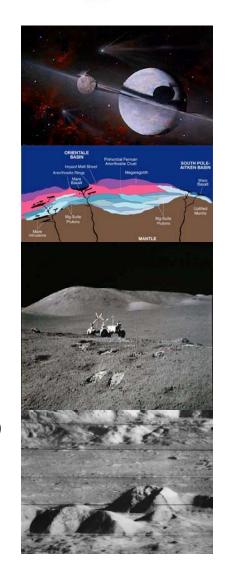
Cumulative New Data for the Moon



- Detailed global elemental and mineralogical information in a spatial context.
- Detailed global topography, gravity, and geodesy.
- Primary understanding of polar environments.
- Detailed imagery (including stereo) at a variety of scales.
- Regolith sounding.
-and more!

What do we hope to learn?

- ORTHUR COLOMADO 26-20 FEBRUARI AND
- What is the early history of the Earth-Moon system?
- How does a small planetary body work?
 [thermal evolution; impact record; etc.]
- What processes have created and concentrated diverse materials on the Moon? Where are they? Why?
- What unimagined mysteries does the Moon hold?
- Where are the most productive places to do "field-work"?



International Opportunities Abound in this New Era of Lunar Exploration

Collaboration

- Release data to community
- Invite participation

Coordination

- Optimize independent activities
- Exchange information for planning
- Cross-calibrate instruments

Cooperation

- Plan joint activities and strategy
- Exchange experiments and personnel









